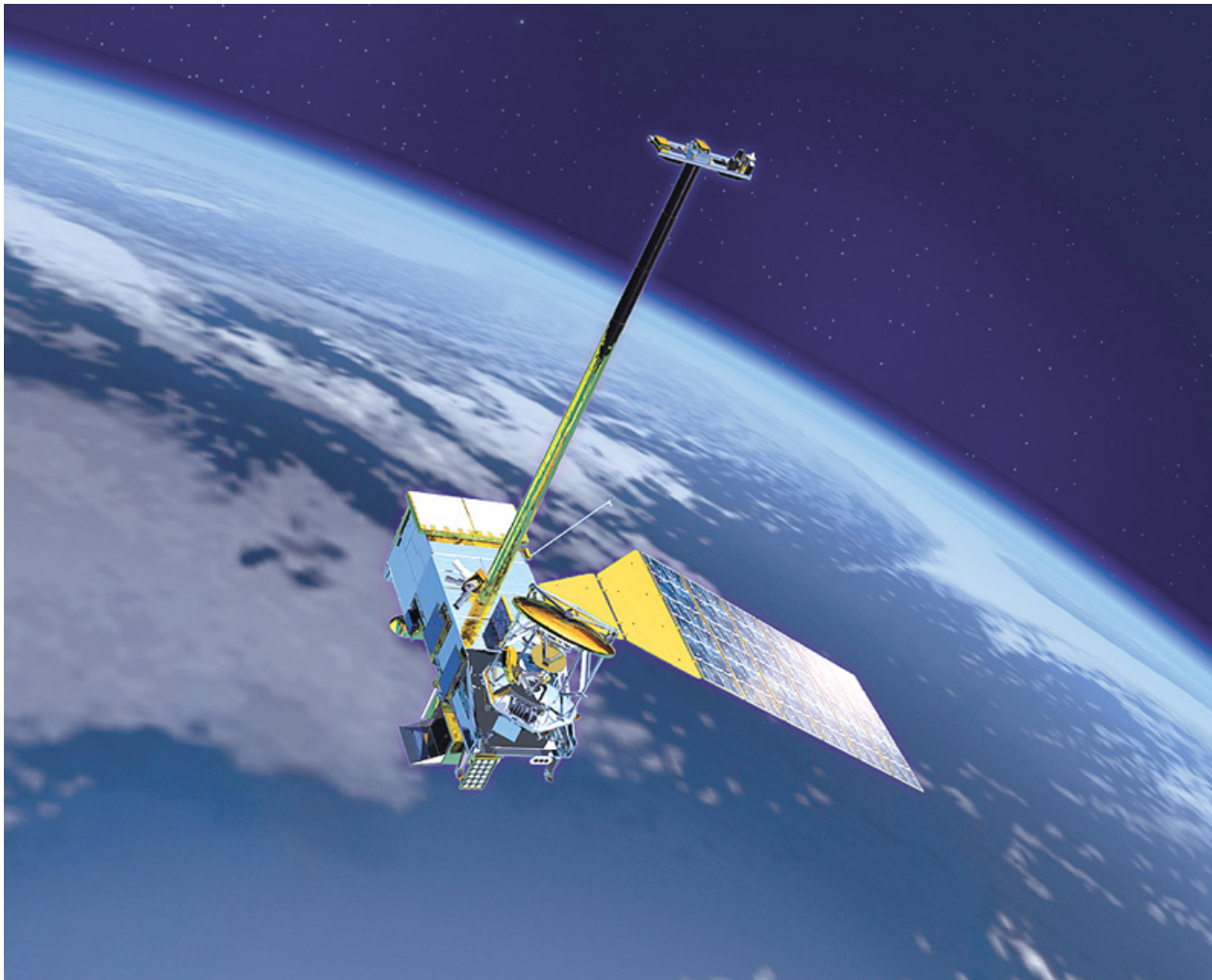


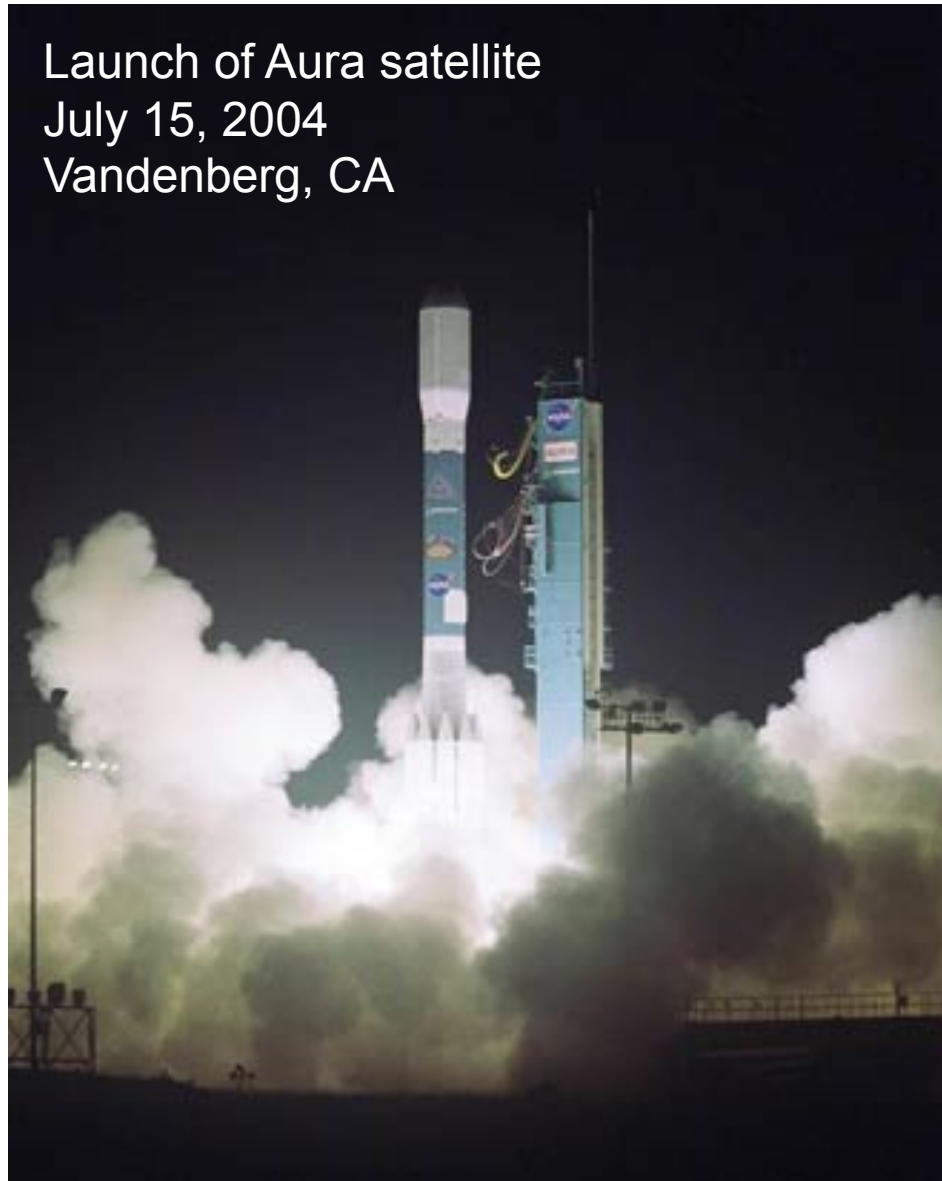
Satellite orbits



Satellite orbits

- Satellites and their orbits
 - GEO
 - LEO
 - Special orbits (e.g., L_1)
- Satellite sensor types
 - Whiskbroom
 - Pushbroom
- Pixel size calculation

Launch



~96% fuel
~4% payload

Recent launch failures:
Orbiting Carbon Observatory (OCO)
Glory

Satellite orbits

- Determined by Kepler's Laws
 - Planets (and satellites) move in elliptical orbits
 - The square of the orbital period is proportional to the cube of the semi-major axis of the orbit
- Altered by atmospheric drag, gravity of Sun and Moon
- Orbital period = time taken for satellite to circle the Earth

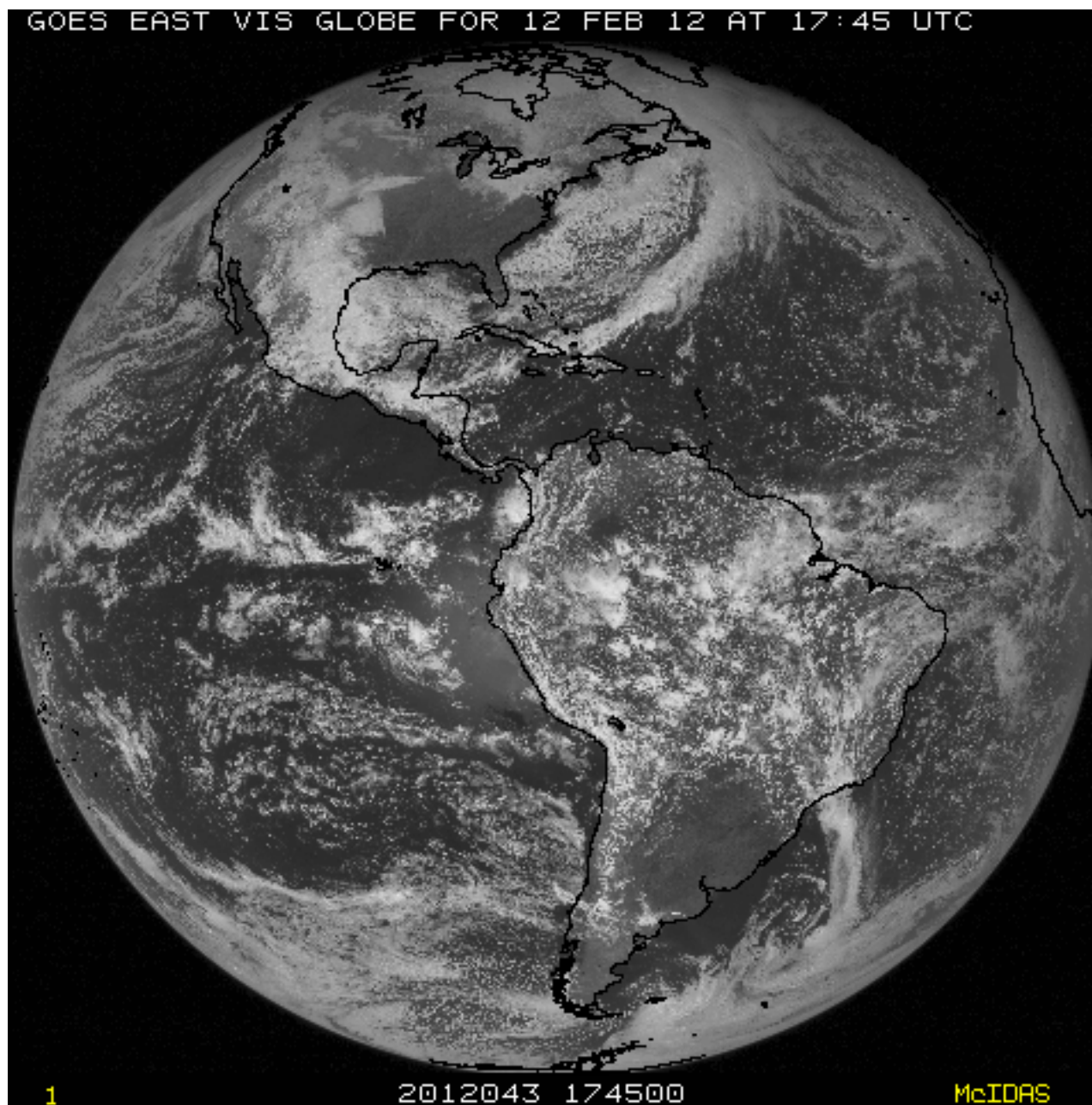
$$P_0 = 2\pi\sqrt{\frac{a^3}{GM}}$$

- P_0 = orbital period, a = semi-major axis of ellipse (satellite altitude + Earth radius [~ 6380 km]), G = Gravitational constant, M = mass of Earth ($GM = 3.986 \times 10^{14} \text{ m}^3 \text{ s}^{-2}$)
- NB. This assumes that the Earth is spherically symmetric (it isn't)

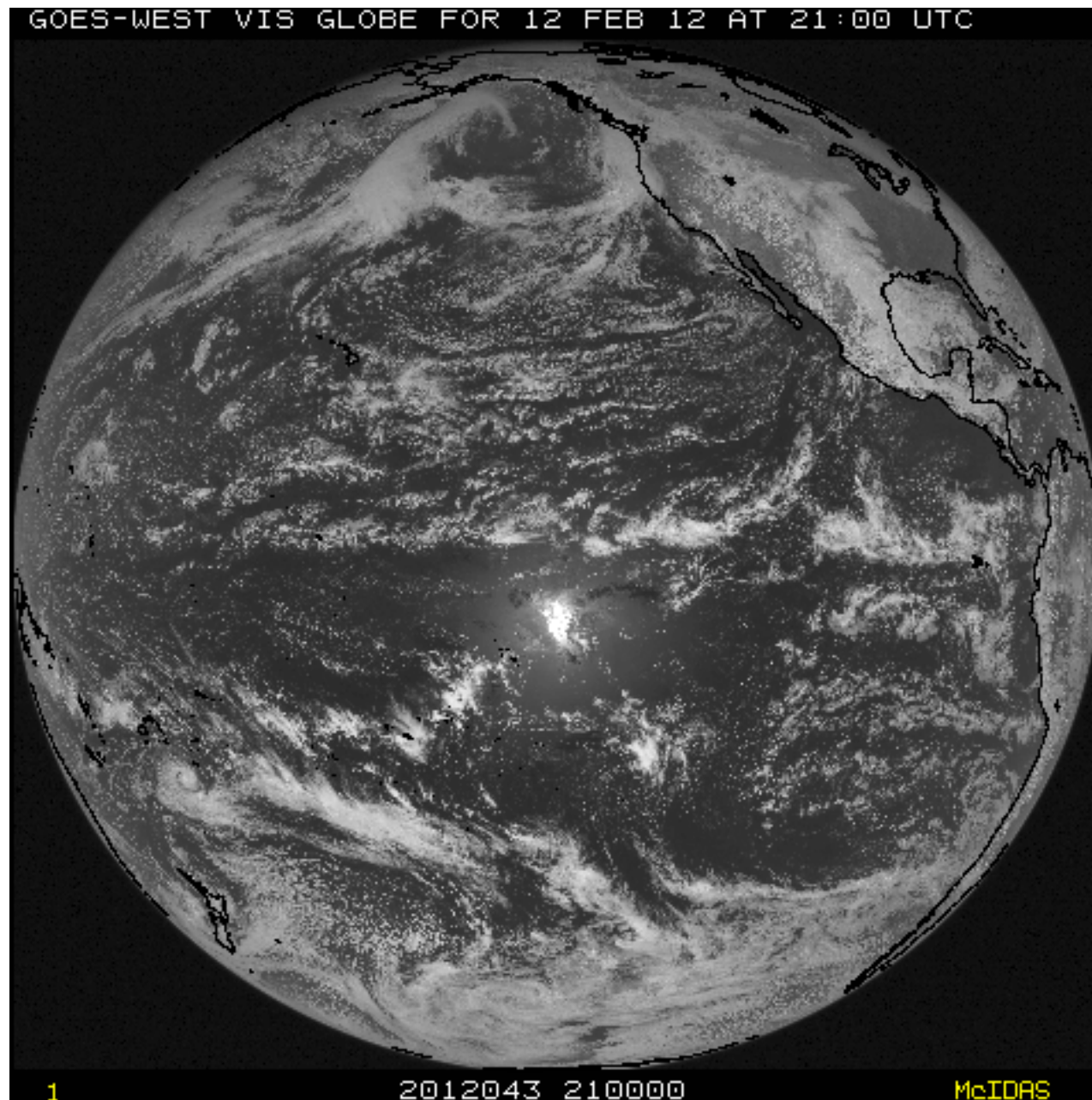
Geostationary orbits (GEO)

- Special type of *geosynchronous* orbit
 - Orbital period = period of Earth's rotation
 - 86164 seconds for a sidereal day
 - Circular orbit around the equator with zero inclination
 - Orbit is stationary with respect to a location on the Earth
-
- Advantages: high temporal resolution, always visible (e.g., weather forecasting, communications); ~5-15 minutes
 - Disadvantages: cost, lower spatial resolution than polar-orbiting sensors, poor coverage of high latitudes ($>55^{\circ}\text{N/S}$)

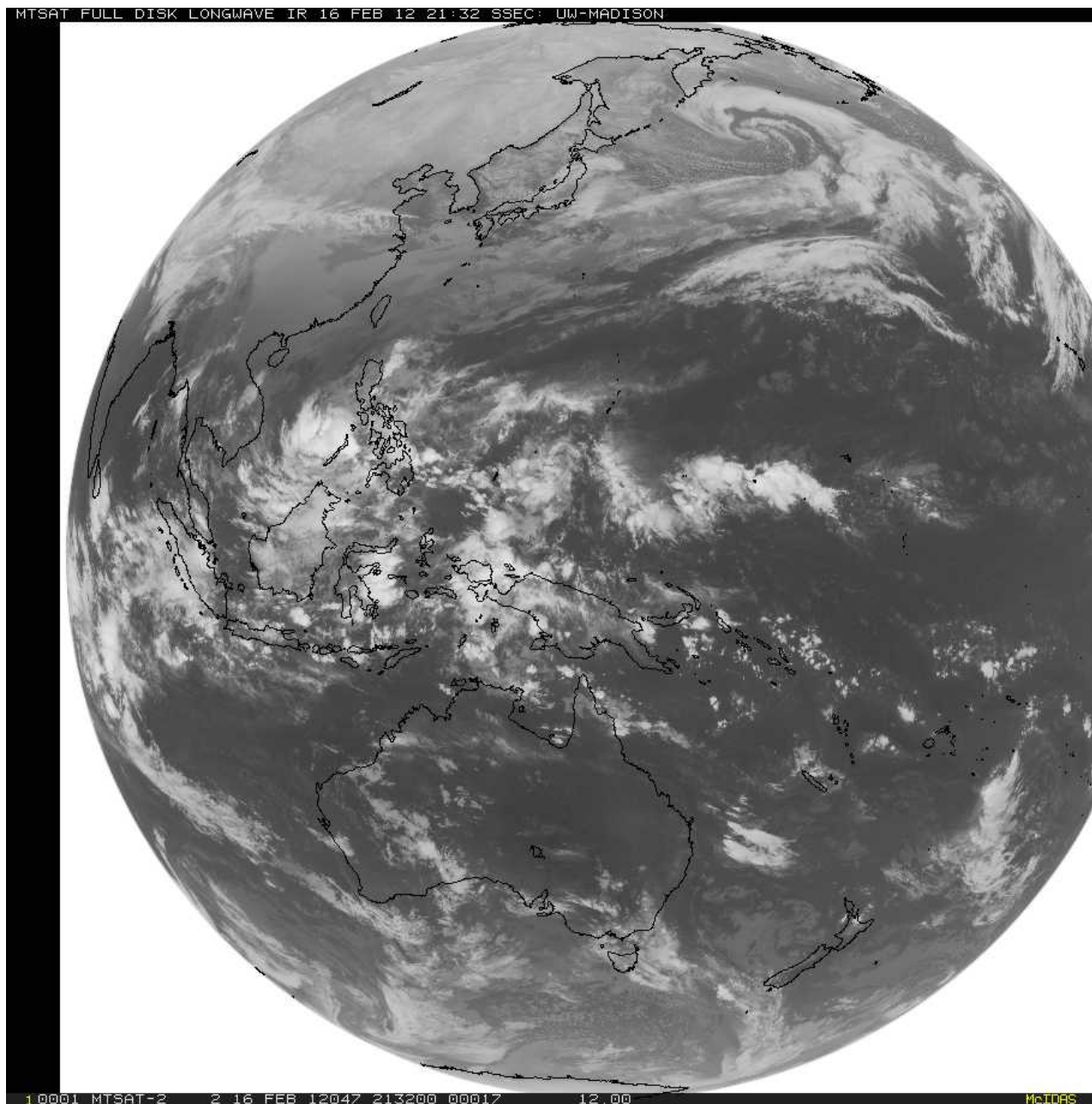
GOES-East (US; 75°W)



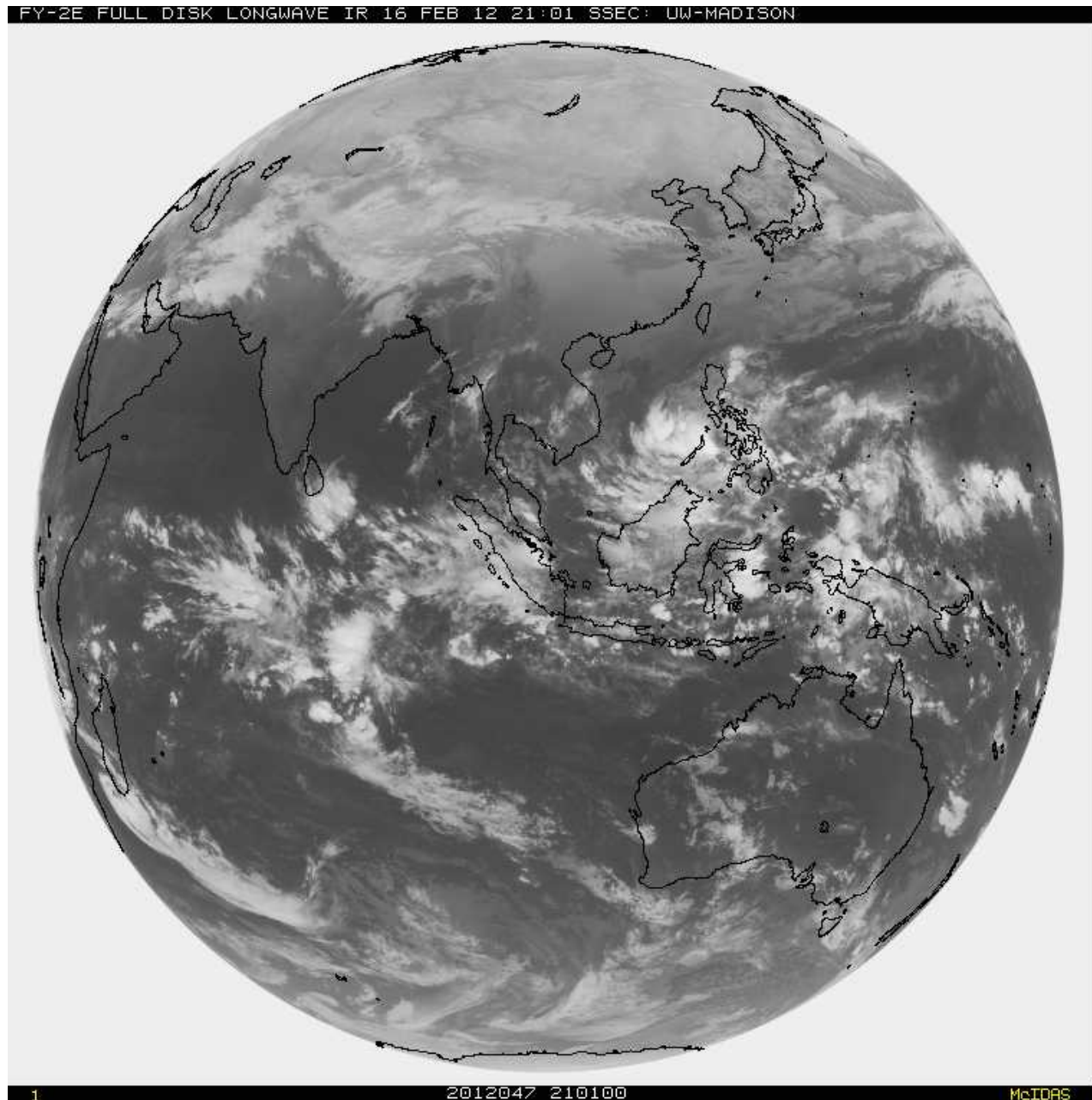
GOES-West (US; 135°W)



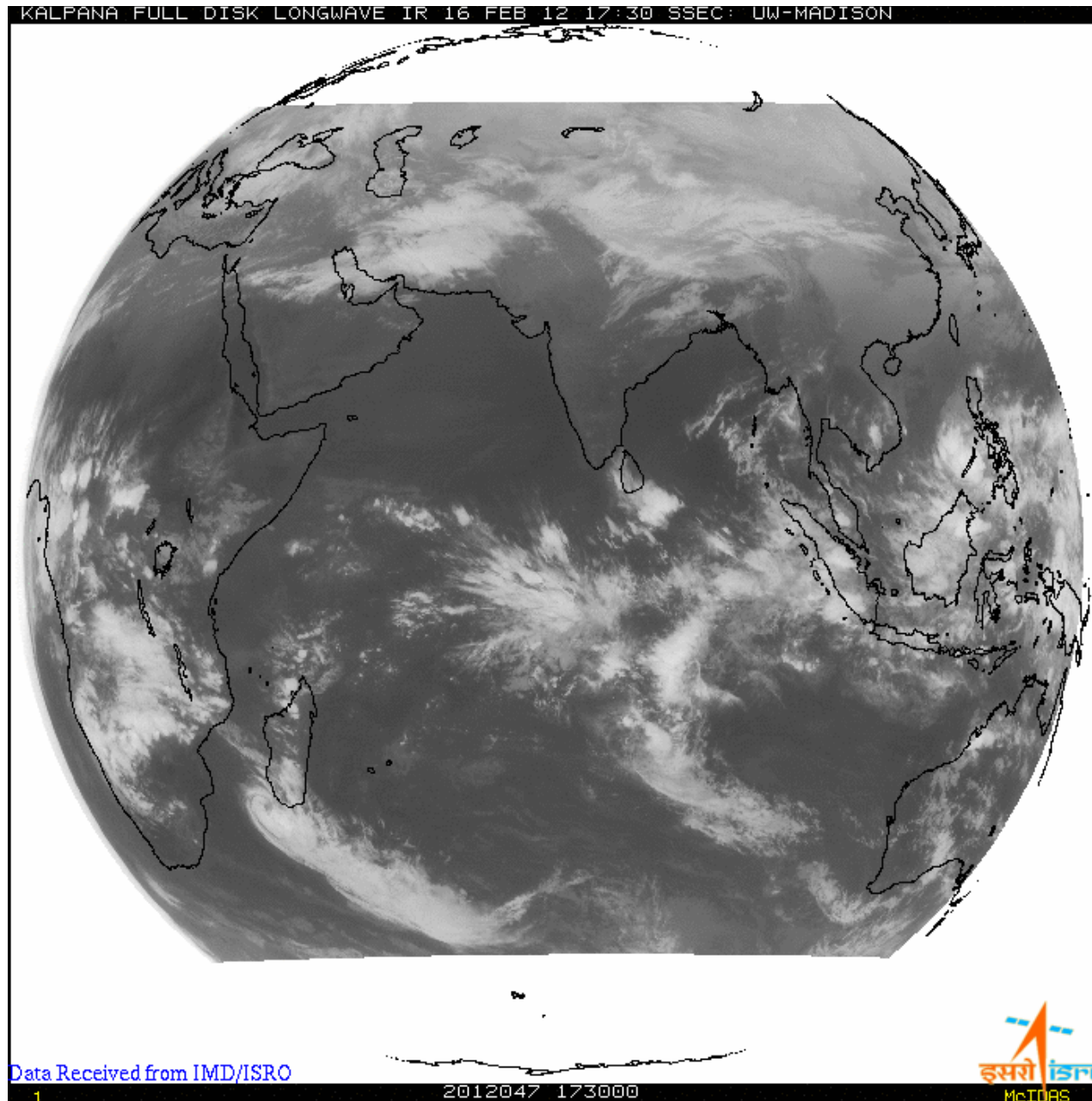
MTSAT (Japan; 140°E)



Fengyun (FY-2E; China; 112°E)

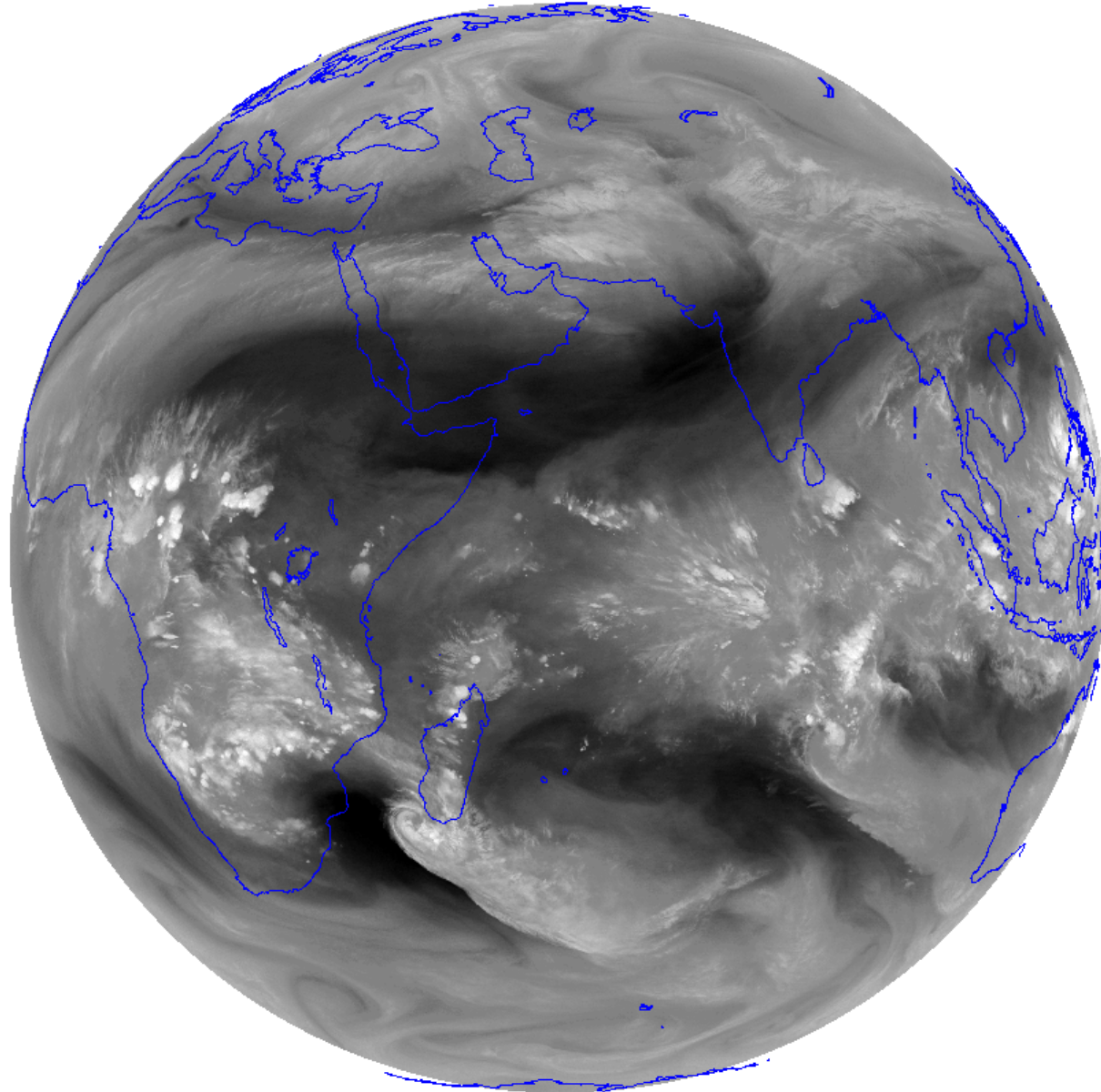


Kalpana (India; 74°E)

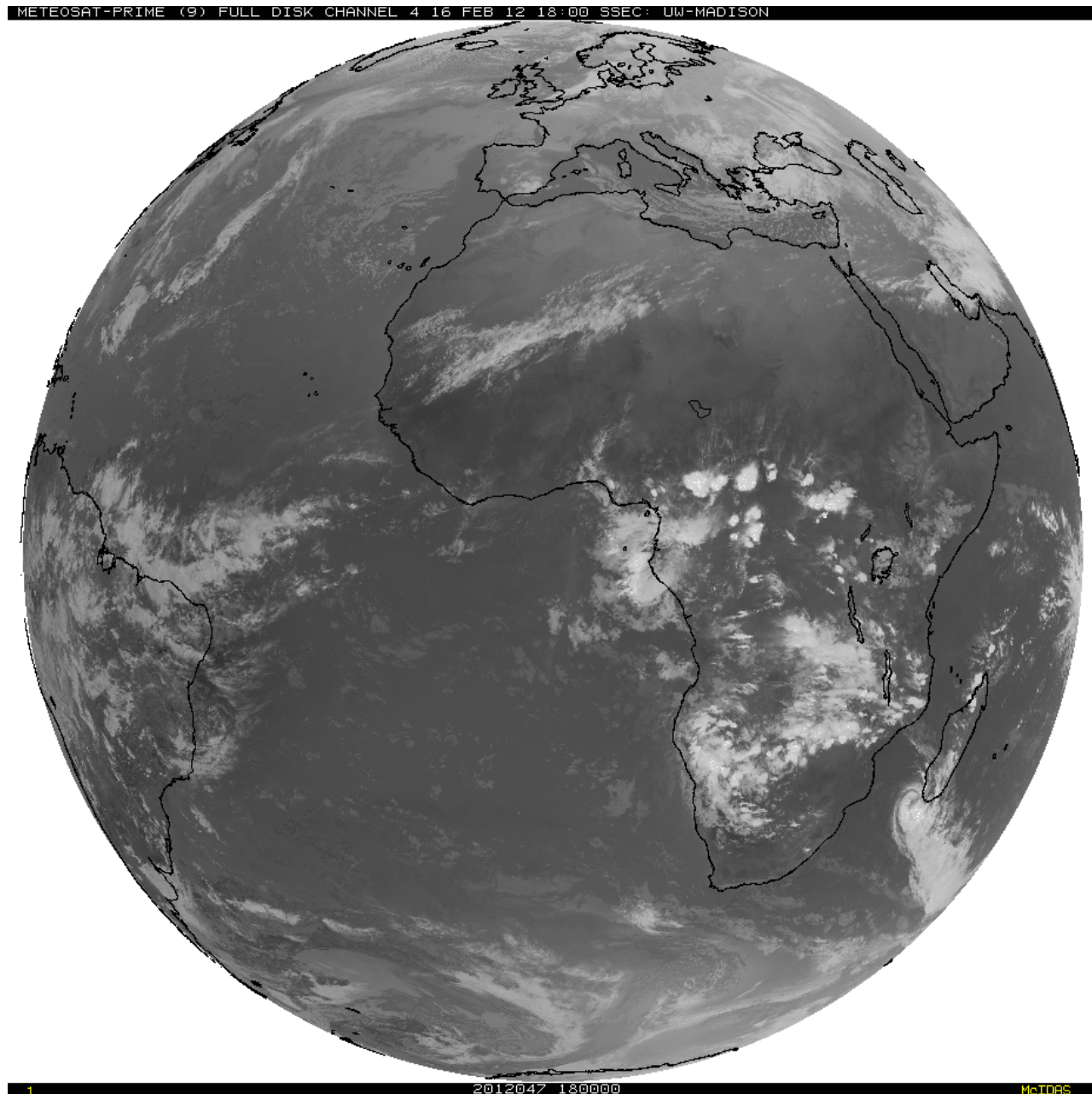


Meteosat-7 (Europe; 57°E)

METEOSAT-10DC (7) FULL DISK WATER VAPOR 16 FEB 12 18:00 SSEC: UW-MADISON



Meteosat-9 (Europe; 0°E)



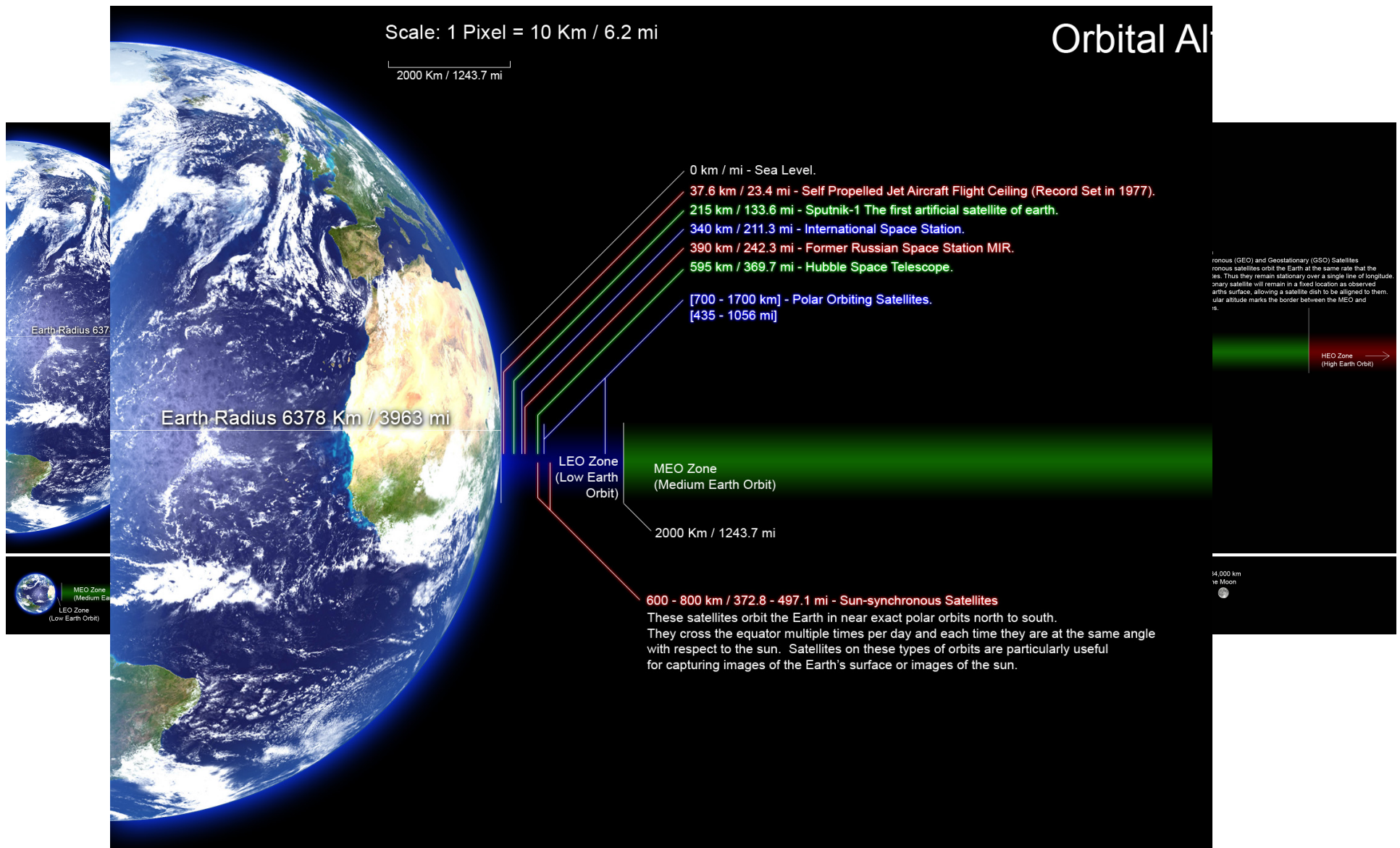
GOES Super Rapid-Scan

A grayscale satellite image from the GOES-15 Imager, showing a wide view of the Pacific Ocean. The image is framed by a blue outline of the Western Hemisphere, including North and South America. The ocean surface is covered with various cloud patterns, including large-scale spiral formations and smaller, more localized cloud clusters. The landmasses of North and South America are visible as darker, more textured areas on the left and right sides of the frame.

<http://cimss.ssec.wisc.edu/goes/blog/archives/6849>

GOES-15 IMAGER - VISIBLE 0.63 (CHANNEL 01) - 20:30 UTC 21 SEPTEMBER 2010 - CIMSS

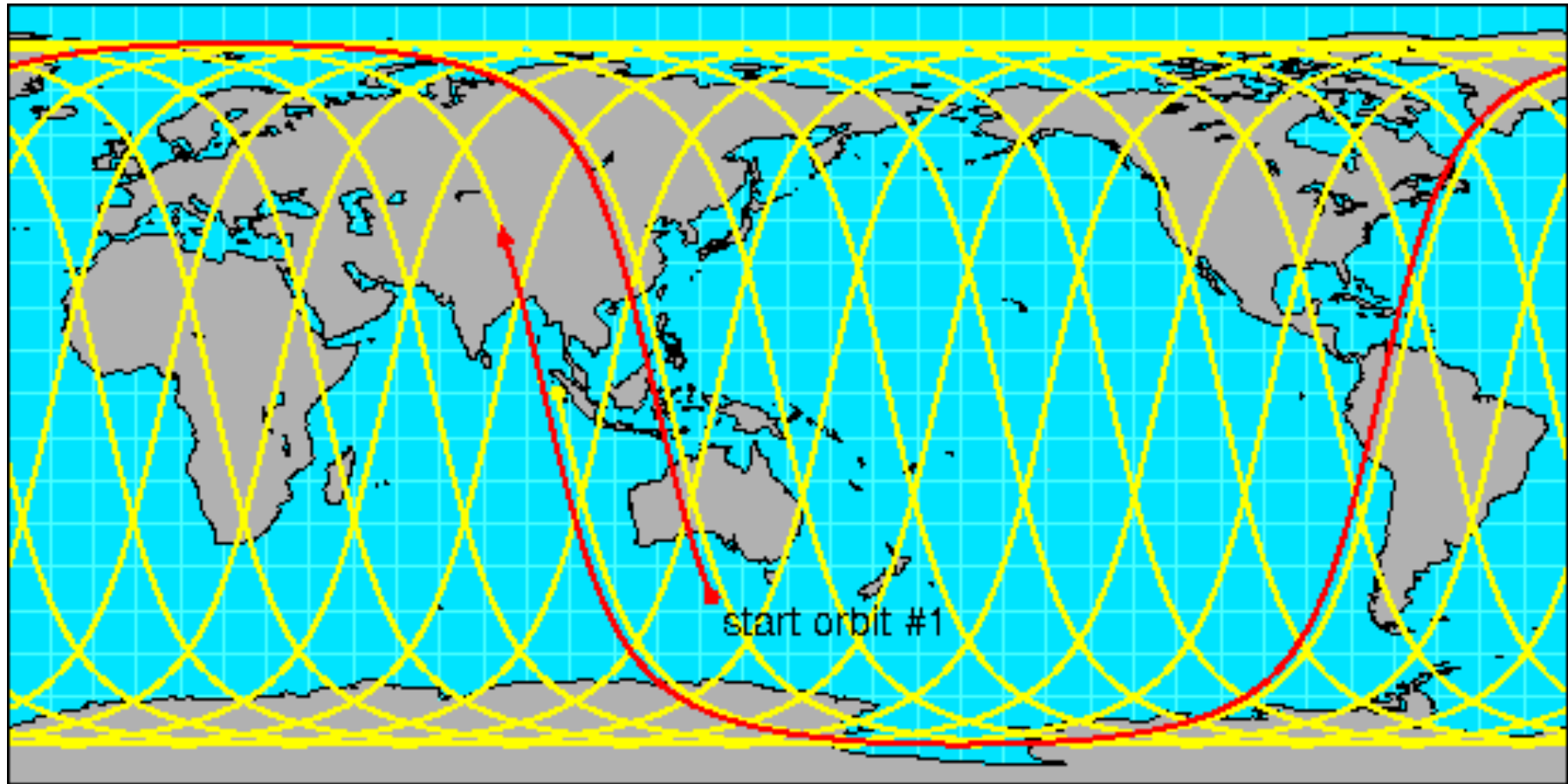
Polar or Low-Earth orbits (LEO)



Polar or Low-Earth orbits (LEO)

- Useful altitude range: ~500-2000 km above Earth's surface
- Constrained by atmospheric friction and van Allen belts (high flux of energetic charged particles)
- Because Earth is not spherical, polar orbits *precess* (rotate) about the Earth's polar axis
- *Sun-synchronous orbits* precess at the same rate that the Earth orbits the Sun
 - Altitudes ~700-800 km, periods of 98-102 minutes
 - 14-15 orbits per day
 - e.g., NOAA-X satellites (US), MetOp (Europe)
- Advantages: high spatial resolution, polar coverage
- Disadvantages: low temporal resolution (at low latitudes)

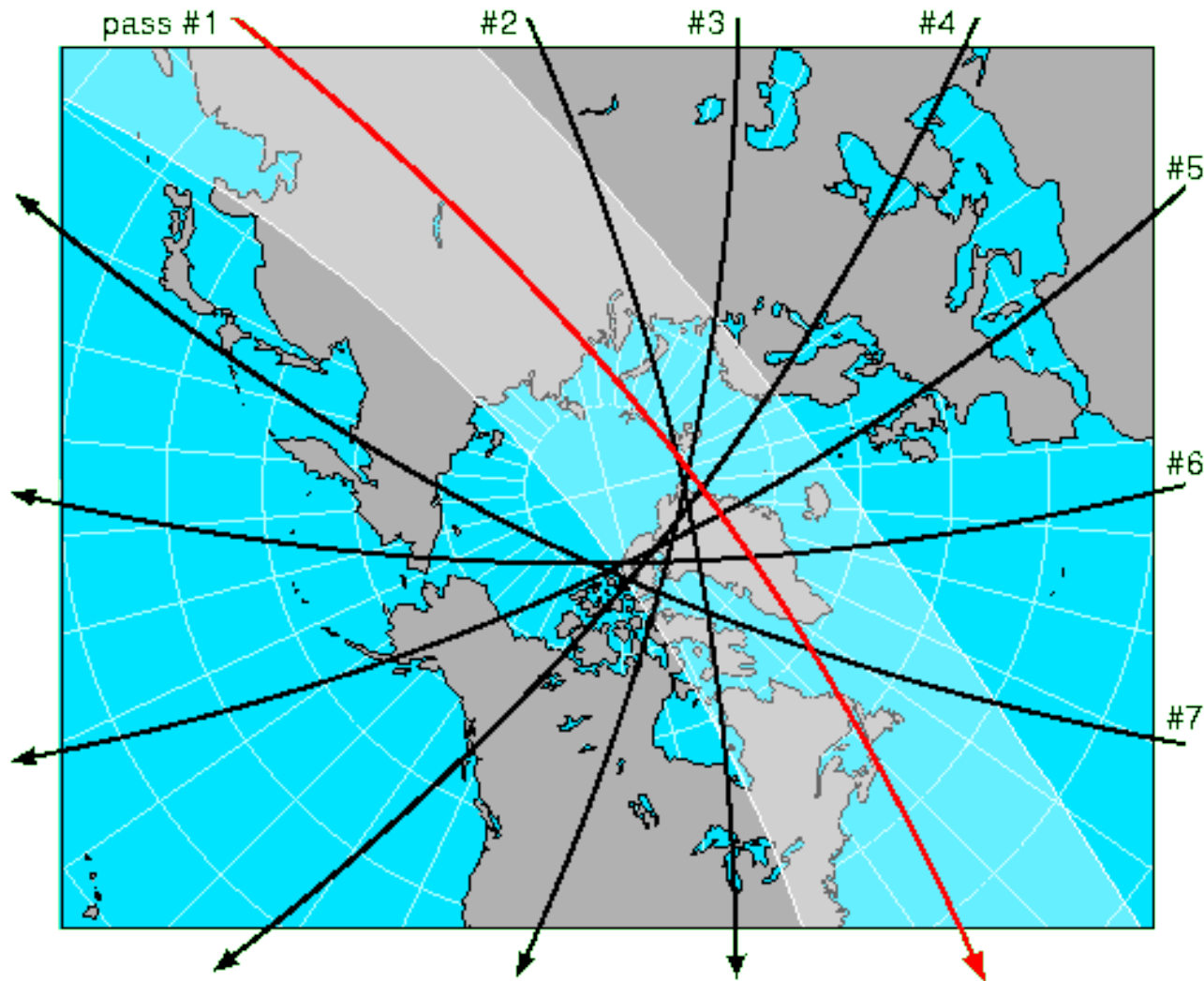
LEO repeat cycles



If Earth makes an integral number of rotations in the time taken for the satellite to complete an integral number of orbits, the sub-satellite track repeats exactly.

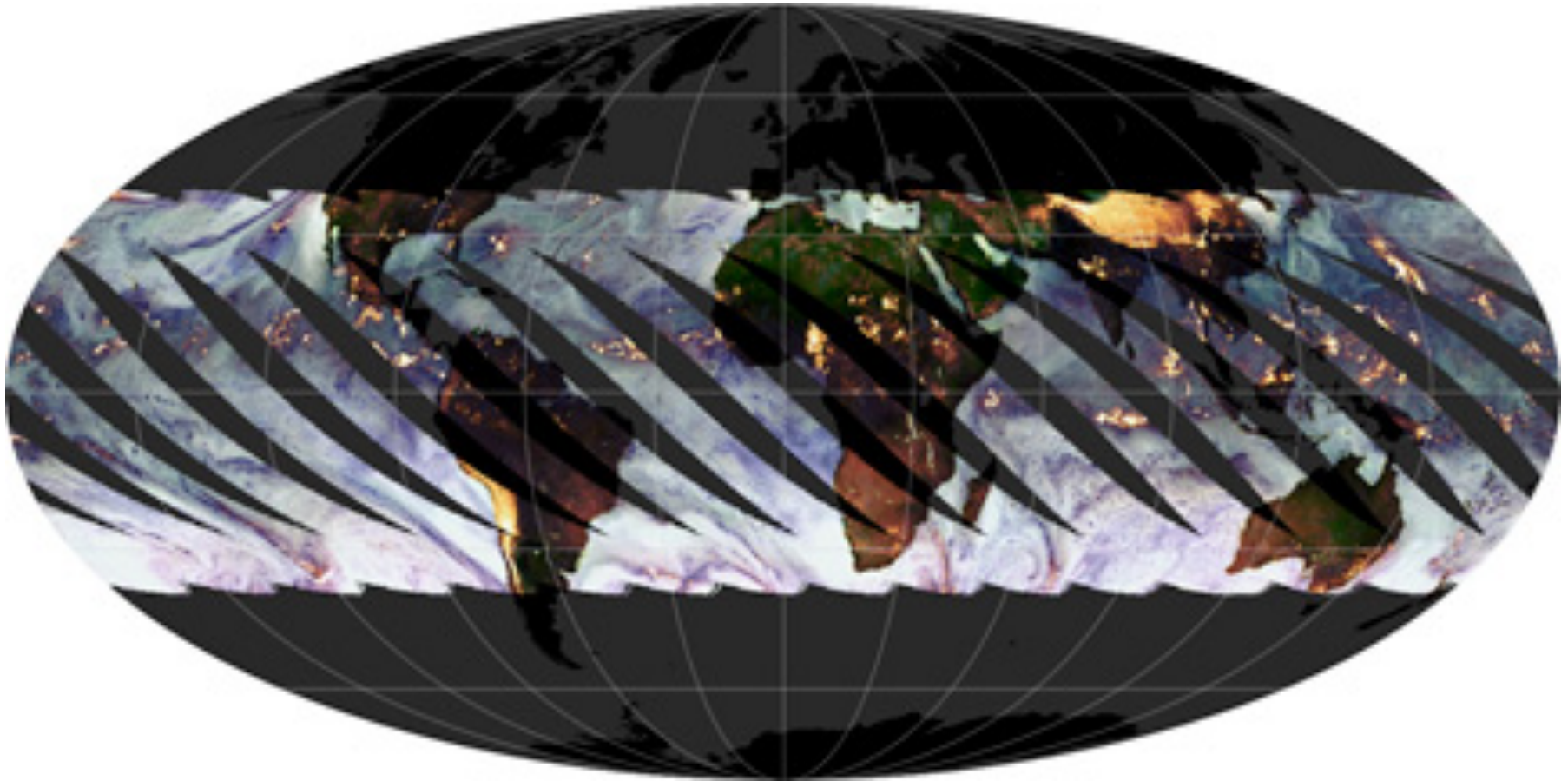
e.g., NASA Aura satellite (705 km altitude) has a 16-day (233 orbit) repeat cycle

LEO polar coverage



LEO orbits converge at the Poles, providing higher temporal resolution

LEO at low inclination

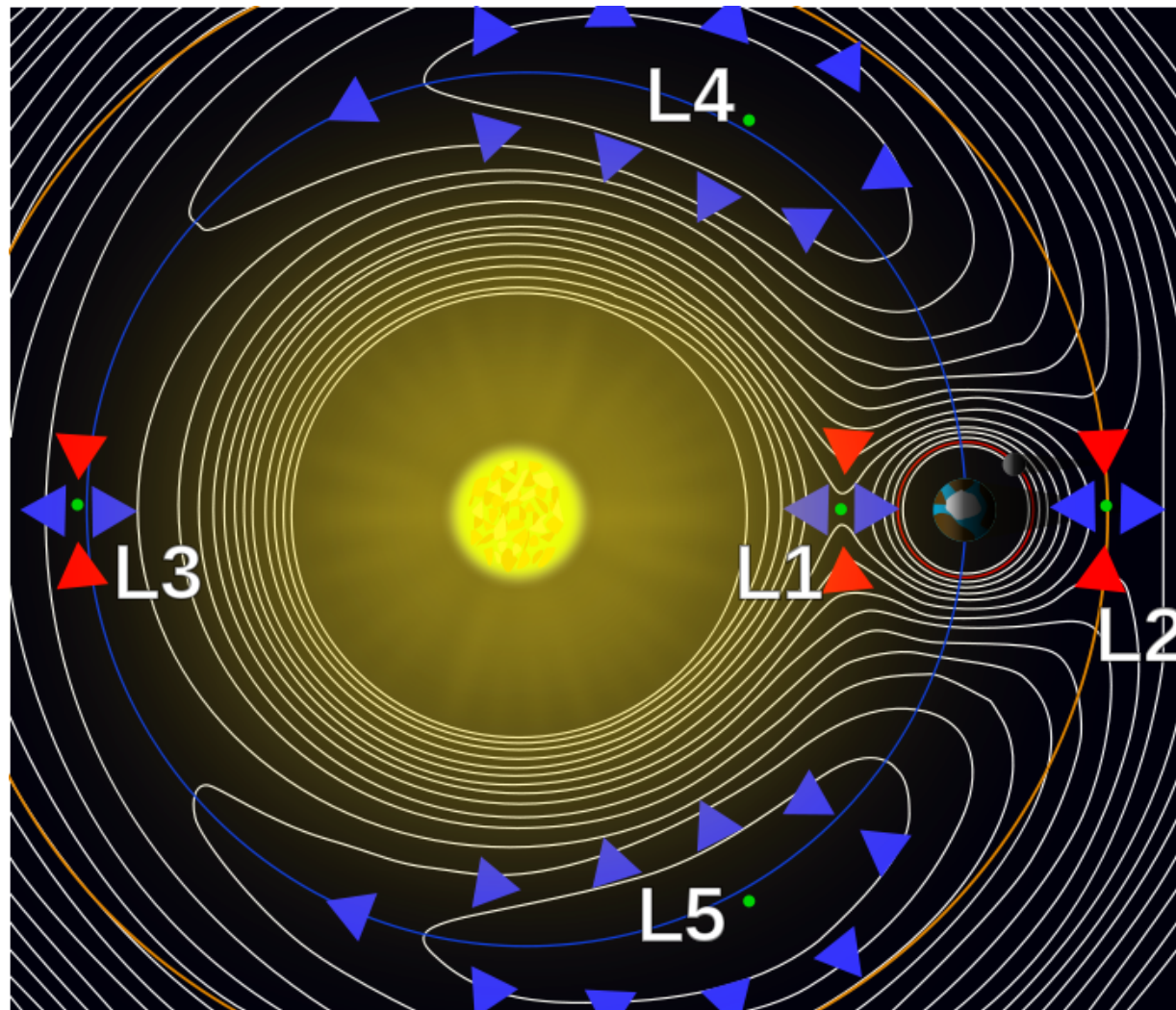


NASA Tropical Rainfall Measuring Mission (TRMM) – 35° inclination, 403 km altitude
<http://trmm.gsfc.nasa.gov>

LEO spacecraft constellations

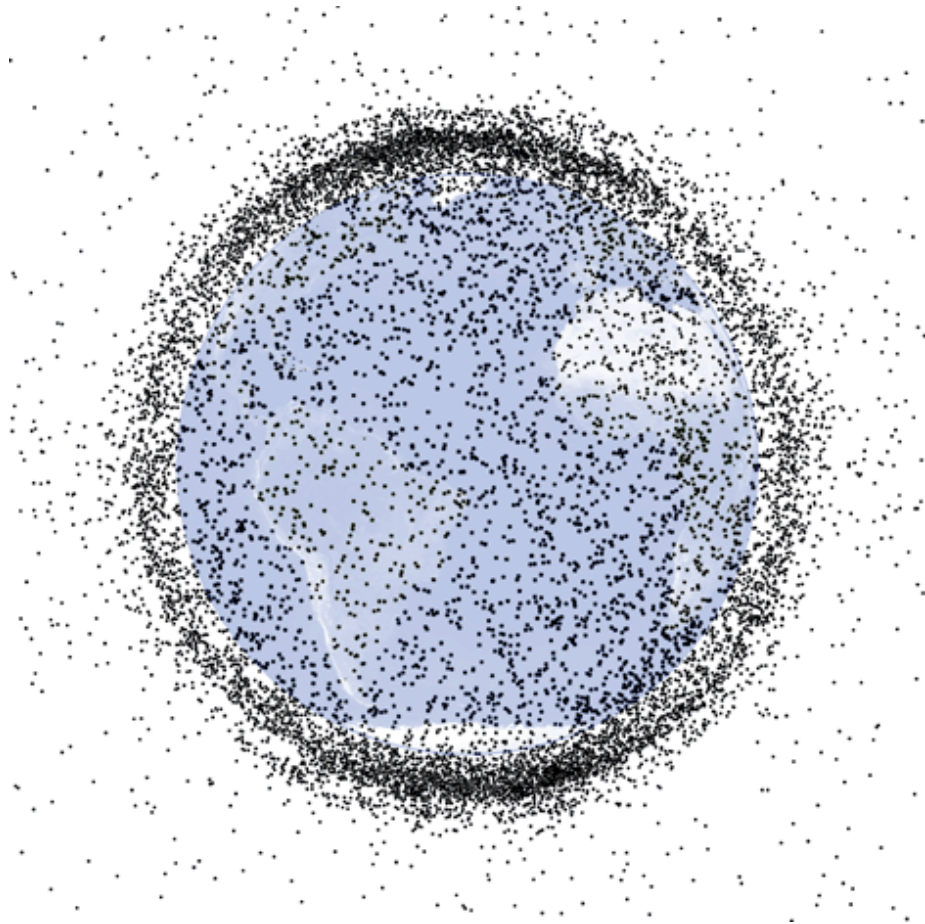


L_1 Lagrange point

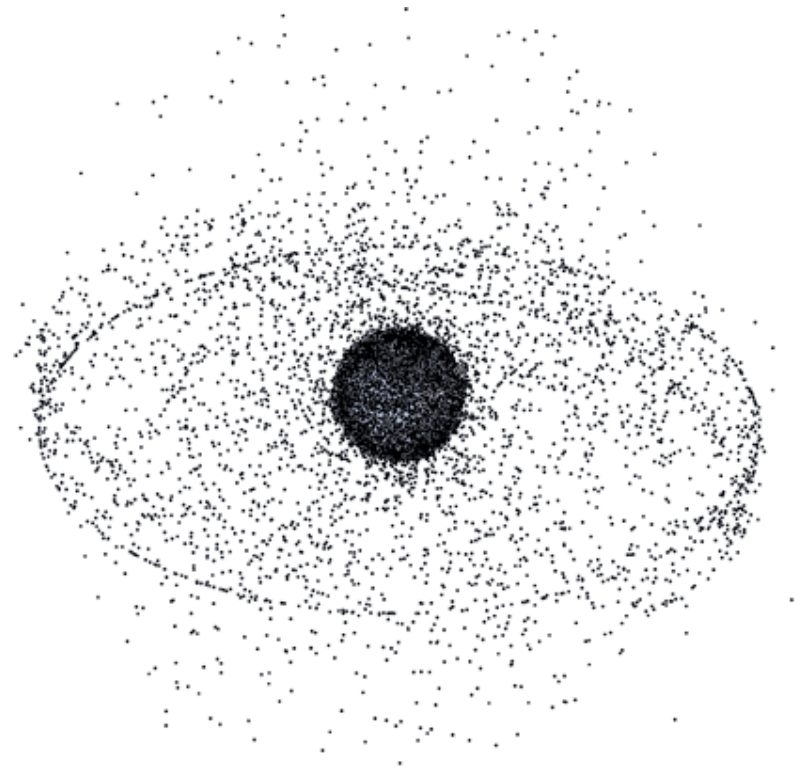


1.5 million kilometers from Earth!

Space junk



LEO



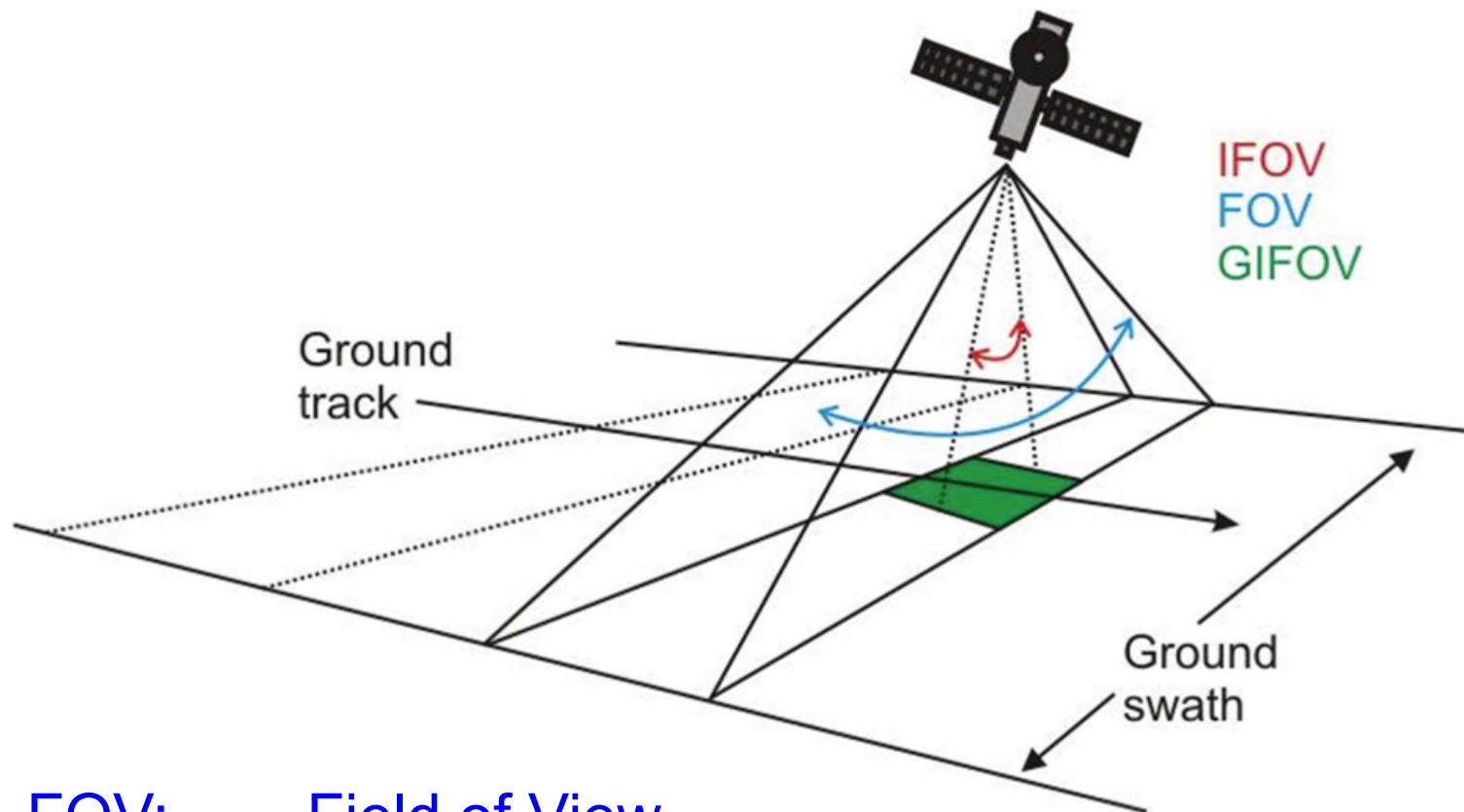
GEO

18,000 manmade objects and counting!

Space-borne imaging systems

- Cross-track scanners
- Whiskbroom scanners
- Pushbroom sensors
- http://www.ssec.wisc.edu/sose/pirs_activity.html

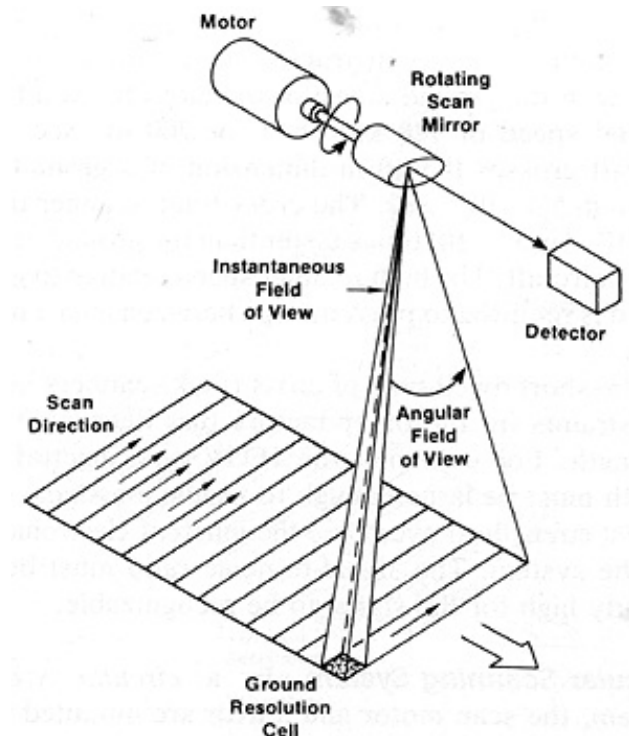
FOV – IFOV - GIFOV



FOV: Field of View
IFOV: Instantaneous FOV
GIFOV: Ground-projected IFOV

Cross-track scanner

- Scans back and forth across the sensor's swath
- Scans each ground-resolution cell (pixel) one-by-one
- Instantaneous field of view (IFOV) of sensor determines pixel size
- Satellite moves along the orbital track as sensor scans across-track
- Divided into 'line' and 'whiskbroom' scanners
- Disadvantages: moving parts, expensive, short pixel 'dwell time', pixel distortion

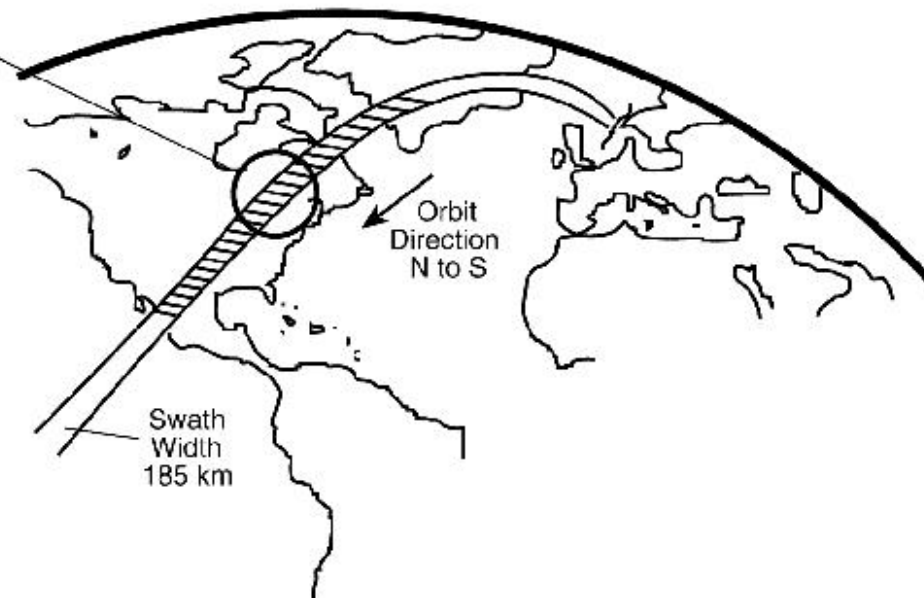
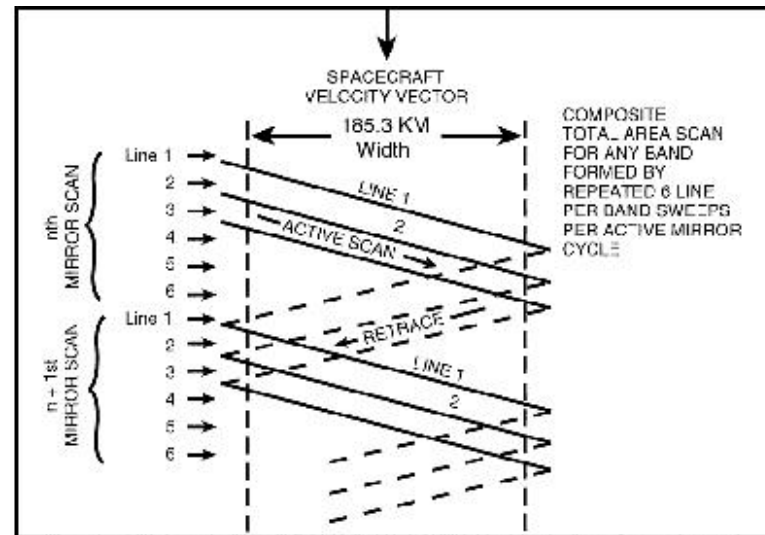
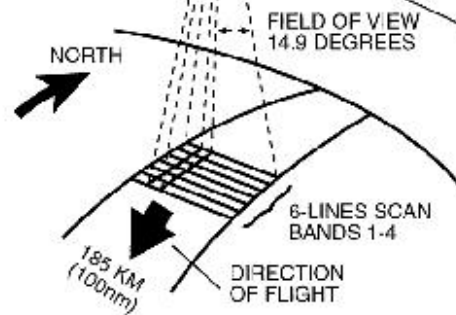
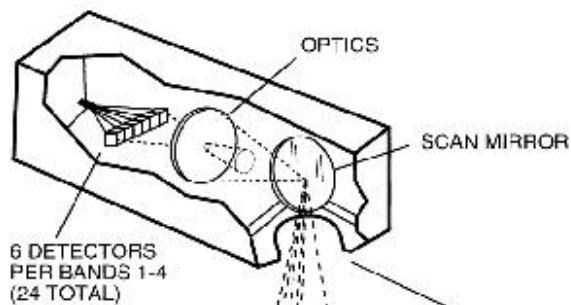


A. CROSS-TRACK SCANNER.

Whiskbroom scanner

Landsat Multi-spectral Scanner (MSS)

MSS Scanning Arrangement



Dwell time (cross-track)

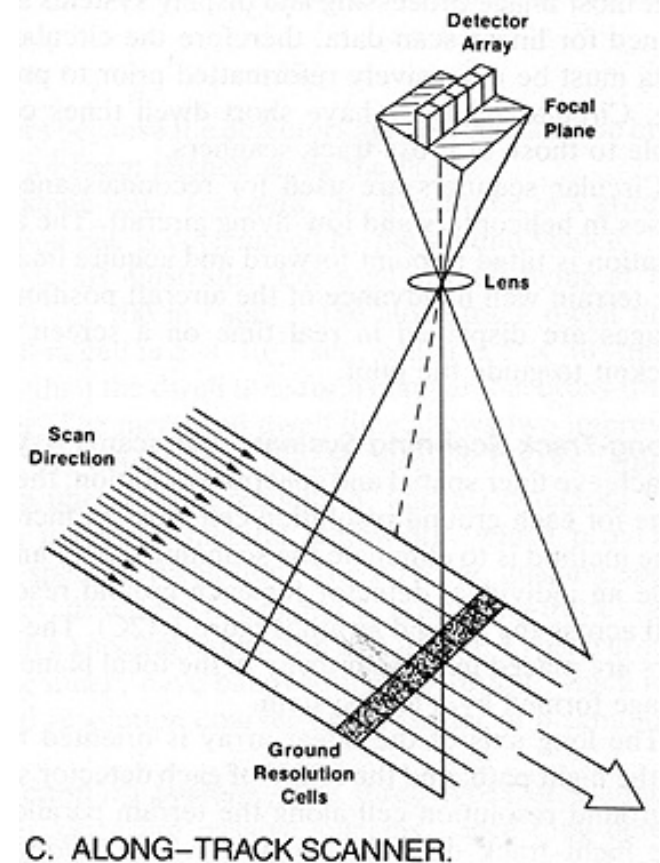
- Time period over which sensor collects photons from an individual ground-resolution cell
- Determines signal-to-noise ratio (SNR)
- Given by [scan time per line]/[# of cells per line]
- Example:

$$\frac{[\text{along-track pixel size}] / [\text{orbital velocity}]}{[\text{swath width}] / [\text{cross-track pixel size}]}$$

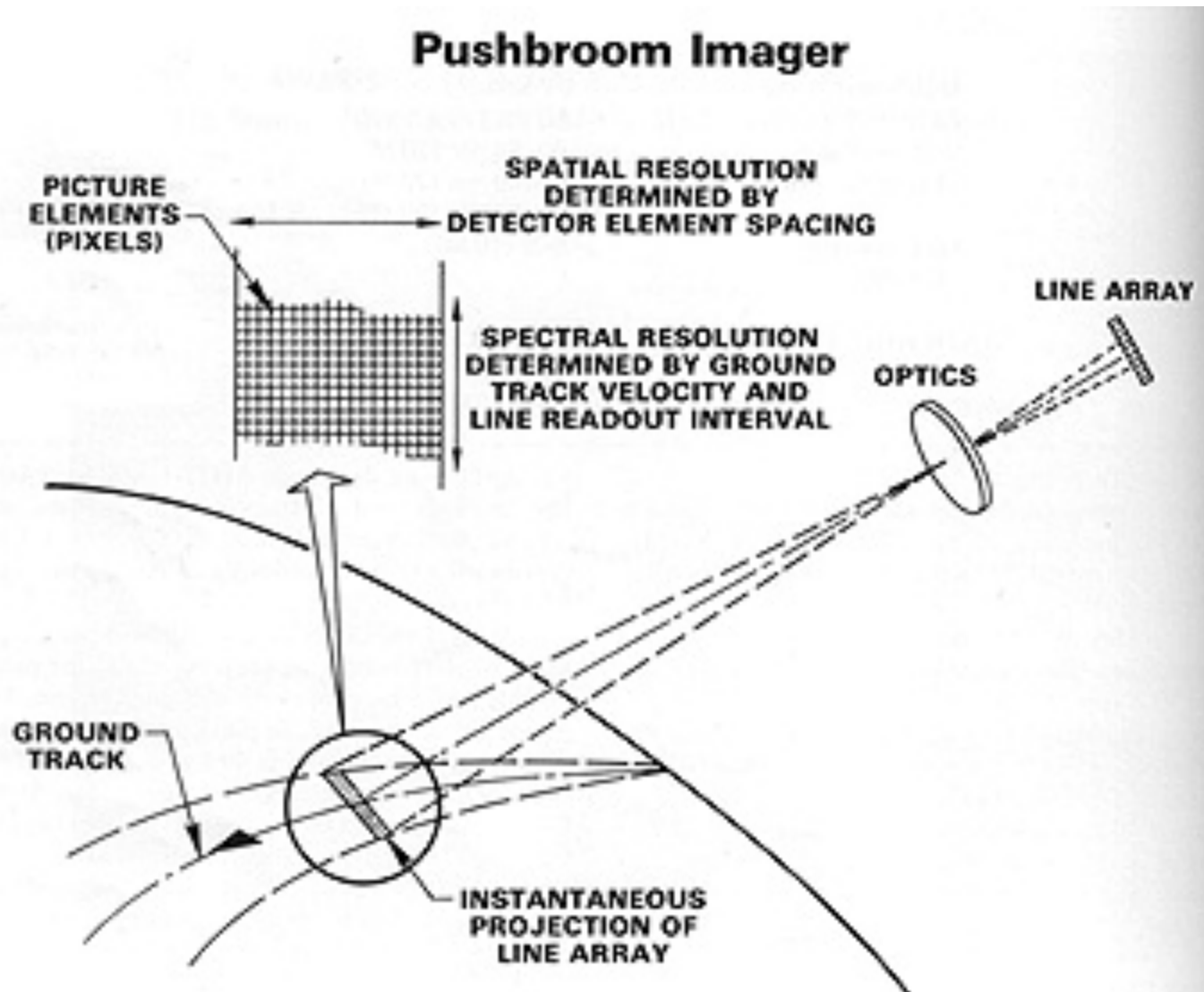
- Landsat sensor with 30×30 m pixel, 185 km swath width; spacecraft velocity = ~7.5 km/s

Along-track scanner (pushbroom)

- Linear array of detectors aligned across-track (e.g., CCD)
- Image built up by satellite movement in flight direction (no scanning mirror)
- 2D detector arrays can acquire multi-spectral or hyperspectral data
 - Optics disperse wavelengths across detector array



Along-track scanner (pushbroom)



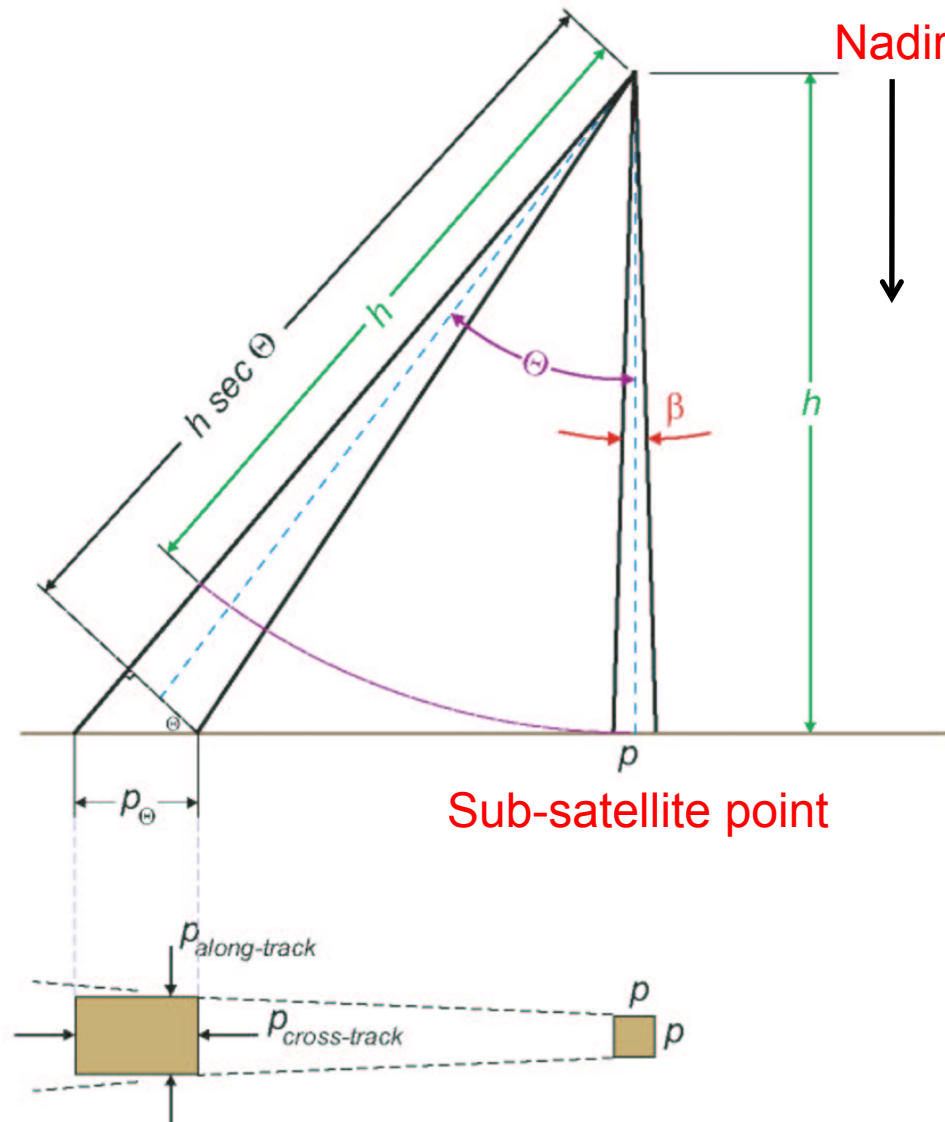
Dwell time (pushbroom)

- Time period over which sensor collects photons from an individual ground-resolution cell
- Determines signal-to-noise ratio (SNR)
- Denominator = 1 in equation below in this case
- Example:

$$\frac{[\text{along-track pixel size}] / [\text{orbital velocity}]}{[\text{swath width}] / [\text{cross-track pixel size}]}$$

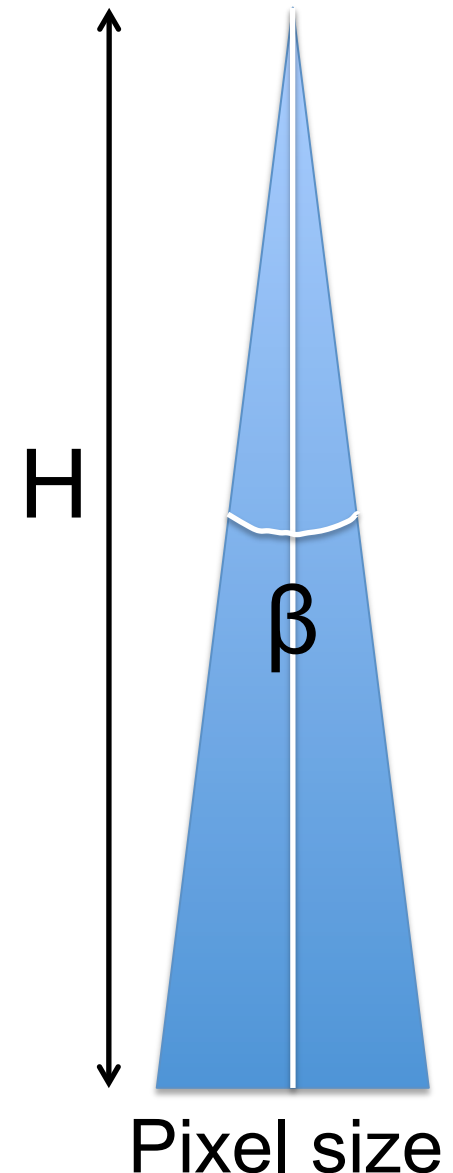
- Different sensitivities and responses in each detector pixel can cause 'striping' in pushbroom sensor data

Satellite viewing geometry



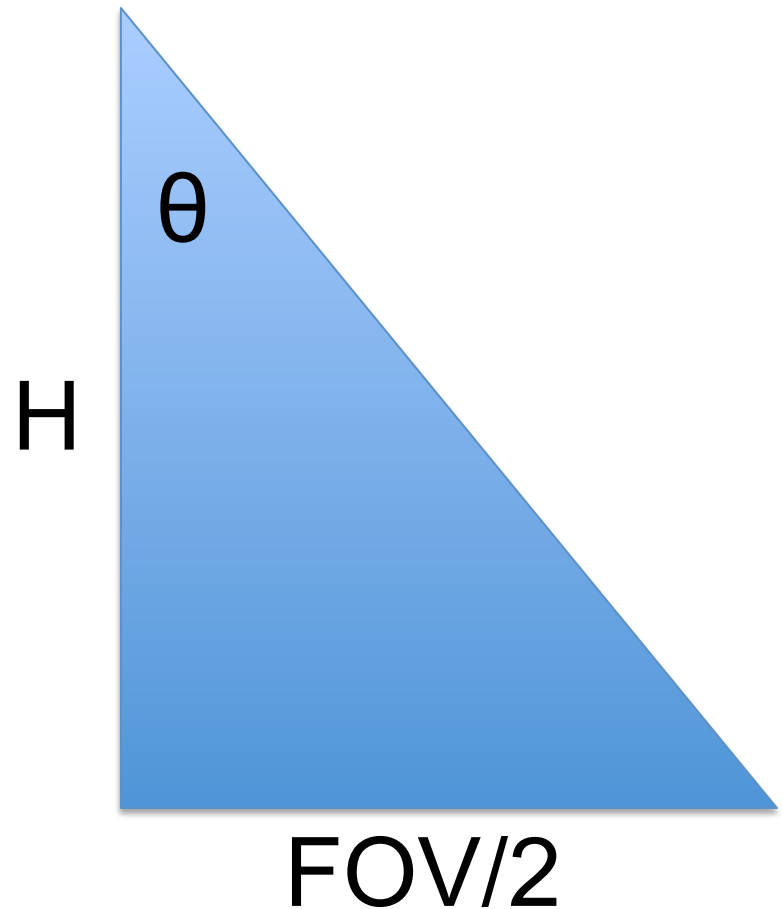
Pixel size calculation

- β = Instantaneous Field of View (IFOV)
- H = satellite altitude
- Pixel size = $2 H \tan (\beta/2)$
- Example:
 - Aura satellite altitude = 705 km
 - OMI (Ozone Monitoring Instrument)
 - OMI telescope IFOV in flight direction = 1°
 - Pixel size = $1410 \tan (0.5) = 12.3$ km
 - *NB: strictly speaking, this is the Ground-projected IFOV (GIFOV) – pixel size could be different*

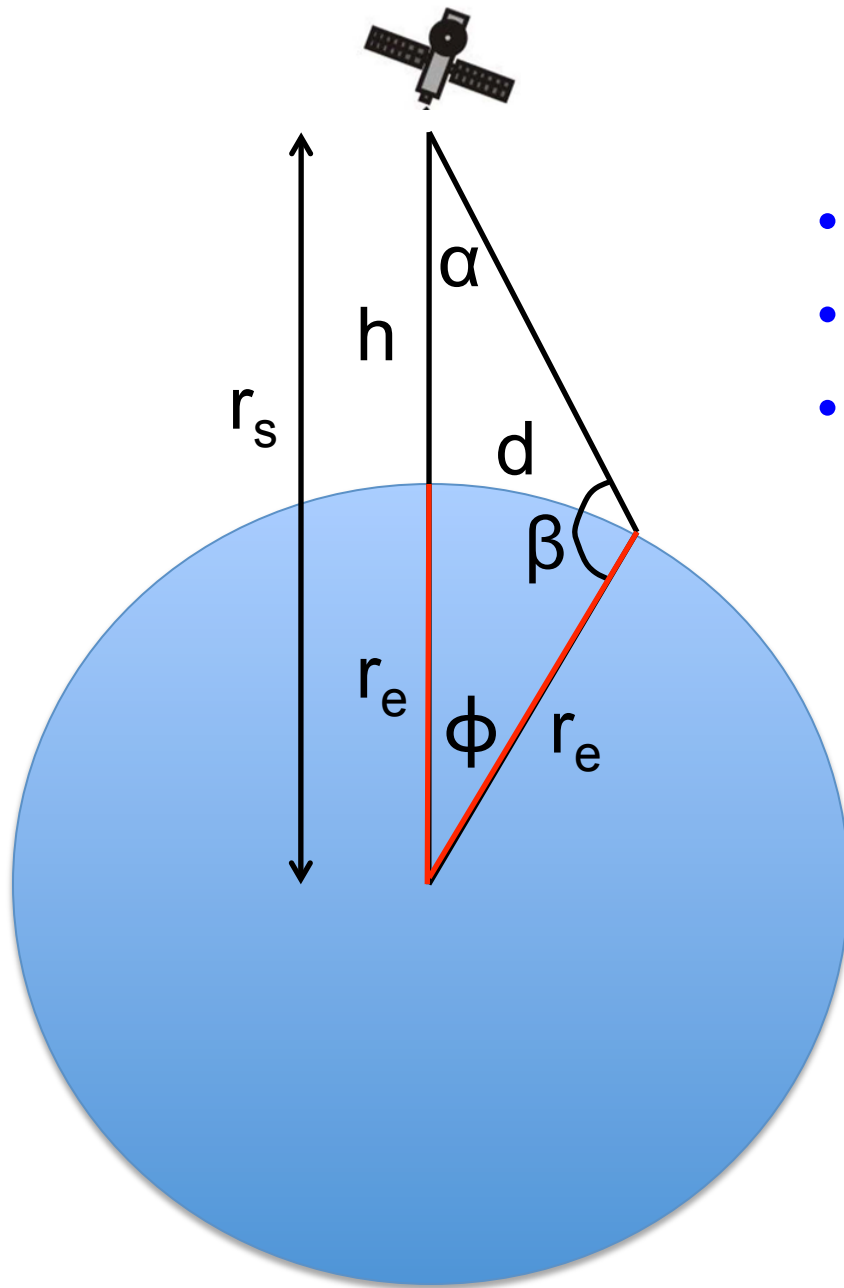


Field of View (FOV)

- $FOV = 2 H \tan (\text{scan angle} + \beta/2)$
- H = satellite altitude
- Example:
 - Aura satellite altitude = 705 km
 - OMI telescope swath FOV = 115°
 - $FOV = 1410 \tan (57.5) = 2213 \text{ km}$
- *But this assumes a flat Earth...*



Field of View (FOV)



- Use Law of Sines
- Note ambiguity for $\sin \beta$
- Swath width = $2d$

$$\sin \beta = \left(\frac{r_s}{r_e} \right) \sin \alpha$$

$$\phi = \pi - \beta - \alpha$$

$$d = \phi r_e$$

Off-nadir viewing geometry

